# Physics \& Astronomy Comprehensive Exam, UCLA, Fall 2012 

 (For those taking the exam for the first time)
## 1. Classical Mechanics (Version 1)

A mass $m_{1}$ moves around a hole on a frictionless horizontal plane. The mass is tied to a massless string of fixed length which passes through the hole. A mass $m_{2}$ is tied to the other end of the string and is subject to uniform gravity with acceleration constant $g$ (see figure below).

(a) Given the initial position $\mathbf{R}_{\mathbf{0}}$ and velocity $\mathbf{V}_{\mathbf{0}}$ in the plane of the table and the masses $m_{1}$ and $m_{2}$, find the equation that determines the maximum and minimum radial distances of the orbit. (Do not attempt to solve this equation.)
(b) Find the frequency of oscillation of the radius of the orbit when the orbit is only slightly different from circular.

## 2. Classical Mechanics (Version 1)

A $K$-meson of mass $m_{K}=494 \mathrm{MeV}$ decays into a $\mu$-meson of mass $m_{\mu}=106 \mathrm{MeV}$ and a neutrino of approximately zero mass $m_{\nu}=0$. Calculate the kinetic energies of the $\mu$-meson and of the neutrino for a $K$-meson decaying while at rest.

## Physics \& Astronomy Comprehensive Exam, UCLA, Fall 2012

(For those repeating the exam)

## 1. Statistical Mechanics (Version 2)

Calculate the lowest non-zero contribution to the specific heat $C_{V}$ of an ideal Fermi gas at low temperatures $T$, as a function of the one-particle density of states $D(\varepsilon)$ at energy $\varepsilon$. The value of the following integral may be helpful,

$$
\int_{-\infty}^{\infty} d x \frac{x^{2}}{\left(e^{x}+e^{-x}\right)^{2}}=\frac{\pi^{2}}{24}
$$

## 2. Electromagnetism (Version 2)

A current sheet of infinite extent along the $y$ and $x$-directions is at a distance $d$ from an infinite planar wall located at $z=0$, as sketched in the diagram below.



The current density of the sheet oscillates in time $t$, and is given by

$$
\vec{j}(x, y, z)=\lambda \cos (\omega t) \delta(z-d) \hat{y} .
$$

The wall exhibits a power reflection coefficient $R$ at frequency $\omega$. The remaining regions are vacuum.
(a) Apply complex notation to find the electric and magnetic field vectors for $z>0$. Use the Cartesian coordinates indicated in the diagram.
(b) Find the time averaged power per unit area required to maintain the current flowing in the sheet.

## 3. Quantum Mechanics

The magnetic dipole moment of the electron is $\boldsymbol{\mu}=-2 \mu_{B} \mathbf{S}$, where $\mu_{B}=\frac{e \hbar}{2 m c}$ is the Bohr magneton ( $e>0$ ), $\mathbf{S}$ the spin.
(a) Write down the Hamiltonian for an electron in a uniform magnetic field $\mathbf{B}$ along the $\mathbf{z -}$ axis (we consider only spin degrees of freedom in this problem, i.e. no momentum).
(b) At $\mathrm{t}=0$, the state of the electron is: $|\psi(0)\rangle=\frac{1}{\sqrt{2}}\binom{1}{1}$ (where the notation is referred to the Pauli matrices). Find $|\psi(t)\rangle$, and $\left\langle\hat{S}_{x}\right\rangle,\left\langle\hat{S}_{y}\right\rangle$ on the state $|\psi(t)\rangle$. Describe your result in physical terms (i.e. what is the spin doing?).
(c) Now introduce a small uniform magnetic field $\mathbf{B}_{1}$ along the $x$-axis. Calculate in perturbation theory the first non-zero correction to the energy levels of the Hamiltonian of part (a). Show that your result is consistent with the exact result (which you can easily calculate).

## 4. Quantum Mechanics

Consider a particle subject to a one-dimensional simple harmonic oscillator potential, whose Hamiltonian is

$$
H=\frac{p^{2}}{2 m}+\frac{k x^{2}}{2}
$$

where $p=-i \hbar \partial_{x}$ is the particle's momentum operator, $m$ mass, and $k$ spring constant. Suppose that at $t=0$ the state vector is given by

$$
e^{-i p a / \hbar}|0\rangle
$$

where $|0\rangle$ is the ground state and $a$ is some number with dimension of length. Evaluate the expectation value of position $\langle x\rangle$ for $\mathrm{t} \geq 0$.

## 5. Quantum Mechanics

A free, spinless non-relativistic particle with mass $m$ and charge $q$ is moving in a uniform magnetic field $\mathbf{B}=\mathrm{Bz}$. Find the spectrum of energy eigenvalues.
(a) Write down the Hamiltonian H in terms of the kinetic momentum $\pi$, including the magnetic field via the minimal substitution $\mathbf{p} \rightarrow \pi=\mathbf{p}-\mathrm{qA}$, where $\mathbf{p}$ is the canonical momentum and $\mathbf{A}$ is the vector potential.
(b) Choose a convenient gauge and write $\mathbf{A}$ in terms of $\mathbf{B}$.
(c) Calculate $\left[\pi_{x}, \pi_{y}\right]$.
(d) Write $\pi_{\mathrm{x}}=i \alpha\left(\mathbf{a}-\mathbf{a}^{\dagger}\right)$ and $\pi_{\mathrm{y}}=\alpha\left(\mathbf{a}^{+} \mathbf{a}^{\dagger}\right)$, where $\alpha$ is a constant, $\mathbf{a}$ is an annihilation operator, and $\mathbf{a}^{\dagger}$ is a creation operator. Find $\alpha$.
(e) Find the eigenvalues of H .

## 6. Quantum Mechanics

A particle in a spherically symmetrical potential is known to be in an eigenstate of $\mathbf{L}^{2}$ and $L_{z}$ with eigenvalues $\hbar^{2} l(l+1)$ and $m \hbar$, respectively, denoted by $|l m\rangle . \mathbf{L}$ is the angular momentum operator, whose components obey the usual commutation algebra. Prove that the expectation values involving $L_{x}$ and $L_{y}$ obey

$$
\left\langle L_{x}\right\rangle=\left\langle L_{y}\right\rangle=0, \quad\left\langle L_{x}^{2}\right\rangle=\left\langle L_{y}^{2}\right\rangle=\frac{l(l+1)-m^{2}}{2} \hbar^{2}
$$

in the eigenstate $|l m\rangle$.

## Questions for the Comprehensive Exam Fall 2012

## 7. Quantum Mechanics

A system that has three unperturbed states can be represented by the perturbed Hamiltonian matrix

$$
\left(\begin{array}{lll}
E_{1} & 0 & a \\
0 & E_{1} & b \\
a^{*} & b^{*} & E_{2}
\end{array}\right) .
$$

The quantities $a$ and $b$ are to be regarded as perturbations that are of the same order and are small compared with $E_{2}-E_{1}$.
(a) Use the second-order nondegenerate perturbation theory to calculate the perturbed eigenvalues.
(b) Diagonalize the matrix to find the exact eigenvalues.
(c) Use the second-order degenerate perturbation theory. Compare the results obtained.

## 8. Statistical Mechanics

This problem concerns the fundamental definitions of thermodynamics. The numbers have been kept small to minimize the mathematics - use discrete differences in place of derivatives where appropriate.
A system has $N=2$ distinguishable particles, each of which can have energy $0, \varepsilon, 2 \varepsilon, 3 \varepsilon, \ldots \infty \varepsilon$. Say the system has total energy $E=5 \varepsilon$.
(a) What is the entropy $S$ ?
(b) What is the temperature $T$ ?
(c) What is the chemical potential $\mu$ ? Determine this by adding a $3^{\text {rd }}$ particle to the system.

Imagine a different system with $N=6$ distinguishable particles, each of which can have energy 0 or $\varepsilon$.
(d) Plot the entropy as a function of the different discrete values of the total $E$.
(e) For what value of $E$ is the temperature maximized?
(f) Plot $T$ as a function of $E$ for the case where $N$ is large (but fixed), and comment on all of the various limiting cases.

## 9. Statistical Mechanics

Consider capillary (surface tension driven) waves at the free surface of a film of liquid He of area A . The dispersion relation is:

$$
\omega=\sqrt{\frac{\sigma}{\rho}} k^{3 / 2} \quad(\sigma \text { is the surface tension and } \rho \text { the density) }
$$

Treating these excitations as quasi-particles obeying Bose statistics, find the contribution of these capillary waves to the heat capacity of the film, specifically, the temperature dependence of this contribution. (Hint: the energy of the quasi-particle is, of course, $E_{\mathrm{k}}=\hbar \omega$.)

## 10. Statistical Mechanics

Consider a collection of $N$ classical and spinless, non-interacting charged particles of charge $q$ and mass $m$ in a region of volume $V$ in which a uniform magnetic field points along the zdirection, i.e., $\vec{B}=B_{0} \hat{z}$, with $B_{0}$ a constant. The charges are in good contact with a heat reservoir at temperature $T$.
(a) Deduce what is the equation of state for this magnetized system.
(b) Find the average induced magnetization $\langle M\rangle$ for this system.
(c) Find the relative (i.e., percentage) magnitude of the RMS fluctuation in average energy $\langle E\rangle$.

## 11. Electromagnetism

Consider an infinitely long cylindrical conductor of diameter D with an infinitely long cylindrical channel cut into it. The channel has a circular cross section with diameter $\mathrm{D} / 3$ and is offset from the axis of the conductor by D/6. The cylinder carries a current I out from the paper plane, uniformly distributed across the solid part of the conductor.

(a) Calculate the magnetic field everywhere on the X-Z plane (i.e. the axis that contains the axis of the two cylinders).
(b) Calculate the magnetic field inside the cylindrical hole.

## 12. Electromagnetism

A pulsar emits bursts of radio waves which are observed from the Earth at two different frequencies, $\omega_{1}$ and $\omega_{2}$. An astronomer notes that the arrival time of the bursts is delayed at the lower frequency, that is, the pulse at $\omega_{1}$ arrives after the pulse at $\omega_{2}$. The delay, $\tau$ is due to dispersion in the interstellar medium. Assume this medium consists of ionized hydrogen (called the dilute plasma), $m_{e}$ is the mass of the electron and $N$ the number of electrons per unit volume.
(a) Find the index of refraction of the dilute plasma.
(b) Find the distance from the pulsar to the Earth.

## 13. Electromagnetism

(a) Consider two positrons in a beam at SLAC. The beam has energy of about $50 \mathrm{GeV}\left(\gamma \approx 10^{5}\right)$. In the beam (rest) frame, they are separated by a distance $d$, and positron $e_{2}^{+}$is traveling directly ahead of $e_{1}^{+}$in the Z-axis, shown in the figure (left) below. Write down $\vec{E}, \vec{B}$, the Lorentz force $\vec{F}$ and the acceleration $\vec{a}$ on $e_{1}^{+}$exerted by $e_{2}^{+}$. Do this in both the rest and laboratory frames. (b) The problem is the same as in part (a) except this time the two positrons are traveling side by side as shown in the figure (right) below.


Two positrons separated by a distance of $d$ travel with a velocity of $v$ in the Z axis.

## 14. Electromagnetism

Consider a dielectric medium of infinite extent in all directions. The medium has a tensor dielectric (at zero frequency) given by

$$
\vec{\varepsilon}=\left(\begin{array}{lcc}
\varepsilon_{x x} & 0 & 0 \\
0 & \varepsilon_{y y} & 0 \\
0 & 0 & \varepsilon_{z z}
\end{array}\right)
$$

with $\varepsilon_{x x}=\varepsilon_{y y} \equiv \varepsilon_{\perp} \neq \varepsilon_{z z}$, and where ( $x, y, z$ ) refer to Cartesian coordinates. A point charge of charge $q$ is placed at the origin of the coordinate system.
(a) Find the magnitude of the electric field at an arbitrary point $(x, y, z)$, i.e., $|\vec{E}|$.
(b) Deduce the polarization charge density $\rho_{p}$ induced on the dielectric at an arbitrary point $(x, y, z)$.
(c) Find the total electrical energy density $u_{E}$ at ( $x, y, z$ ).

## Question <br> 1 <br> VERSION

## Question 1

A mass $m_{1}$ moves around a hole on a frictionless horizontal plane. The mass is tied to a massless string of fixed length which passes through the hole. A mass $m_{2}$ is tied to the other end of the string and is subject to uniform gravity with acceleration constant $g$ (see fig 1).
(a) Given the initial position $\mathbf{R}_{0}$ and velocity $\mathbf{V}_{0}$ in the plane of the table and the masses $m_{1}$ and $m_{2}$, find the equation that determines the maximum and minimum radial distances of the orbit. (Do not attempt to solve this equation.)
(b) Find the frequency of oscillation of the radius of the orbit when the orbit is only slightly different from circular.


## Solution to Question 1

The system is conservative and subject to holonomic constraints; thus we use the Lagrangian formulation of mechanics to derive the equations of motion. We use polar coordinates $(r, \theta)$ in the horizontal plane to parametrize the position of $m_{1}$ (and thus of $m_{2}$ ).
(a) The Lagrangian $L$ for the combined system is given by,

$$
\begin{equation*}
L=\frac{1}{2} m_{1}\left(\dot{r}^{2}+r^{2} \dot{\theta}^{2}\right)+\frac{1}{2} m_{2} \dot{r}^{2}-m_{2} g r \tag{0.1}
\end{equation*}
$$

The canonical momenta are,

$$
\begin{align*}
p_{r} & =\left(m_{1}+m_{2}\right) \dot{r} \\
p_{\theta} & =m_{1} r^{2} \dot{\theta} \tag{0.2}
\end{align*}
$$

Since $L$ is independent of $\theta$, the angular momentum $p_{\theta}$ is conserved during the motion. The Euler-Lagrange equation for $r$ is given by,

$$
\begin{equation*}
\left(m_{1}+m_{2}\right) \ddot{r}-m_{1} r \dot{\theta}^{2}+m_{2} g=0 \tag{0.3}
\end{equation*}
$$

Eliminating $\dot{\theta}$, we find a reduced equation for $r$ only,

$$
\begin{equation*}
\left(m_{1}+m_{2}\right) \ddot{r}-\frac{p_{\theta}^{2}}{m_{1} r^{3}}+m_{2} g=0 \tag{0.4}
\end{equation*}
$$

By multiplying through by $\dot{r}$, we integrate the resulting equation to get the total energy $\varepsilon$,

$$
\begin{equation*}
\frac{1}{2}\left(m_{1}+m_{2}\right) \dot{r}^{2}+\frac{p_{\theta}^{2}}{2 m_{1} r^{2}}+m_{2} g r=\varepsilon \tag{0.5}
\end{equation*}
$$

The constants of motion $p_{\theta}$ and $\varepsilon$ are determined by the initial conditions, which we cast in the following form,

$$
\begin{array}{ll}
r=R_{0} & \dot{r}=V_{0} \cos \phi \\
& r \dot{\theta}=V_{0} \sin \phi \tag{0.6}
\end{array}
$$

so that

$$
\begin{align*}
p_{\theta} & =m_{1} R_{0} V_{0} \sin \phi \\
\varepsilon & =\frac{1}{2} m_{1} V_{0}^{2}+\frac{1}{2} m_{2} V_{0}^{2} \cos ^{2} \phi+m_{2} g R_{0} \tag{0.7}
\end{align*}
$$

The condition for extremal radius $r=r_{e}$ is given by setting $\dot{r}=0$, and we find,

$$
\begin{equation*}
\frac{p_{\theta}^{2}}{2 m_{1} r_{e}^{2}}+m_{2} g r_{e}=\varepsilon \tag{0.8}
\end{equation*}
$$

This equation is equivalent to a 3 -rd order polynomial equation which always has one physically unacceptable negative root, and two positive roots representing the minimum and the maximum radius.
(b) Circular motion corresponds to $r=R_{0}$ and $\phi=\pi / 2$. Small deviations from circular motion may be parametrized by $r=R_{0}+x$ with $|x| \ll R_{0}$, and $\phi-\pi / 2$ small. Since the corrections to $p_{\theta}$ due to the perturbation in $\phi$ are second order, we may neglect those. Thus, the fluctuation equation becomes,

$$
\begin{equation*}
\left(m_{1}+m_{2}\right) \ddot{x}+\frac{3 p_{\theta}^{2}}{m_{1} R_{0}^{4}} x=0 \tag{0.9}
\end{equation*}
$$

Using the equilibrium condition for circular motion, $p_{\theta}^{2}=m_{1} m_{2} R_{0}^{3} g$, we find the following frequency $\omega$ of small oscillations,

$$
\begin{equation*}
\omega^{2}=\frac{3 m_{2} g}{\left(m_{1}+m_{2}\right) R_{0}} \tag{0.10}
\end{equation*}
$$

$Q_{2}, V_{1}$

## MOA

## Question 2

A $K$-meson of mass $m_{K}=494 \mathrm{MeV}$ decays into a $\mu$-meson of mass $m_{\mu}=106 \mathrm{MeV}$ and a neutrino of approximately zero mass $m_{\nu}=0$. Calculate the kinetic energies of the $\mu$-meson and of the neutrino for a $K$-meson decaying while at rest.

## Solution to Question 2

Conservation of momentum implies that the momentum $\mathbf{p}_{\mu}$ of the $\mu$-meson and the momentum $\mathbf{p}_{\nu}$ of the neutrino are opposite to one another $\mathbf{p}_{\mu}=-\mathbf{p}_{\nu}$. Thus the energy of the neutrino is given by $E_{\nu}=\left|\mathbf{p}_{\nu}\right| c=\left|\mathbf{p}_{\mu}\right| c$. Conservation of energy reads,

$$
\begin{equation*}
m_{K} c^{2}=\left|\mathbf{p}_{\mu}\right| c+\sqrt{\mathbf{p}_{\mu}^{2} c^{2}+m_{\mu}^{2} c^{4}} \tag{0.11}
\end{equation*}
$$

Solving for $\left|\mathbf{p}_{\mu}\right|$, we find,

$$
\begin{equation*}
\left|\mathbf{p}_{\mu}\right|=\frac{m_{K}^{2}-m_{\mu}^{2}}{2 m_{K}} c \tag{0.12}
\end{equation*}
$$

The kinetic energy $T_{\nu}$ of the neutrino and of the kinetic energy $T_{\mu}$ of the $\mu$-meson are,

$$
\begin{align*}
T_{\nu} & =\frac{m_{K}^{2}-m_{\mu}^{2}}{2 m_{K}} c^{2} \\
T_{\mu} & =\frac{m_{K}^{2}+m_{\mu}^{2}}{2 m_{K}} c^{2}-m_{\mu} c^{2} \tag{0.13}
\end{align*}
$$

Numerical evaluation gives approximate values,

$$
\begin{equation*}
T_{\nu}=236 \mathrm{MeV} \quad T_{\mu}=152 \mathrm{MeV} \tag{0.14}
\end{equation*}
$$

## Question 1, VERSION 2

Calculate the lowest non-zero contribution to the specific heat $C_{V}$ of an ideal Fermi gas at low temperatures $T$, as a function of the one-particle density of states $D(\varepsilon)$ at energy $\varepsilon$. The value of the following integral may be helpful,

$$
\begin{equation*}
\int_{-\infty}^{\infty} d x \frac{x^{2}}{\left(e^{x}+e^{-x}\right)^{2}}=\frac{\pi^{2}}{24} \tag{0.15}
\end{equation*}
$$

## Solution to Question 3

In terms of the density of one-particle states $D(\varepsilon)$ and the Fermi occupation number,

$$
\begin{equation*}
f(\varepsilon)=\frac{1}{e^{\beta(\varepsilon-\mu)}+1} \quad \beta=\frac{1}{k T} \tag{0.16}
\end{equation*}
$$

the total number of particles $N$ and the internal energy $E$ are given by,

$$
\begin{align*}
& N=\int_{0}^{\infty} d \varepsilon D(\varepsilon) f(\varepsilon) \\
& E=\int_{0}^{\infty} d \varepsilon \varepsilon D(\varepsilon) f(\varepsilon) \tag{0.17}
\end{align*}
$$

(as usual, we have assumed that $D(\varepsilon)=0$ for $\varepsilon<0$ ). We also define the chemical potential $\mu_{0}$ at zero temperature by the relation,

$$
\begin{equation*}
N=\int_{0}^{\mu_{0}} d \varepsilon D(\varepsilon) \tag{0.18}
\end{equation*}
$$

Since the function $D(\varepsilon)$ does not involve temperature, the specific heat is given by,

$$
\begin{equation*}
C_{V}=\frac{\partial E}{\partial T}=\int_{0}^{\infty} d \varepsilon \varepsilon D(\varepsilon) \frac{\partial f(\varepsilon)}{\partial T} \tag{0.19}
\end{equation*}
$$

Working this out, and after some minor simplifications, we get,

$$
\begin{equation*}
C_{V}=\frac{1}{k T^{2}} \int_{0}^{\infty} d \varepsilon \frac{\varepsilon(\varepsilon-\mu) D(\varepsilon)}{\left(e^{\beta(\varepsilon-\mu) / 2}+e^{-\beta(\varepsilon-\mu) / 2}\right)^{2}} \tag{0.20}
\end{equation*}
$$

For low temperatures, the denominator is responsible for concentrating the support of the integral over $\varepsilon$ near $\mu$, so we may extend the integration region all the way to $-\infty$. Also, to leading order, we may evaluate $D(\varepsilon)$ at the central value $\mu$. Finally, the parity of the
remaining integral allows us to replace the factor $\varepsilon$ in the numerator by $\varepsilon-\mu$, so that we end up with the following expression,

$$
\begin{equation*}
C_{V}=\frac{D(\mu)}{k T^{2}} \int_{-\infty}^{\infty} d \varepsilon \frac{(\varepsilon-\mu)^{2}}{\left(e^{\beta(\varepsilon-\mu) / 2}+e^{-\beta(\varepsilon-\mu) / 2}\right)^{2}} \tag{0.21}
\end{equation*}
$$

Changing variables from $\varepsilon$ to $x$ with $\varepsilon=\mu+2 k T x$ gives,

$$
\begin{equation*}
C_{V}=8 k^{2} T D(\mu) \int_{-\infty}^{\infty} d x \frac{x^{2}}{\left(e^{x}+e^{-x}\right)^{2}} \tag{0.22}
\end{equation*}
$$

Finally, using the value of the integral stated in the problem, and using the fact that for small temperatures we have $\mu=\mu_{0}+\mathcal{O}\left(T^{2}\right)$, we may approximate this result further by setting $D(\mu)=D\left(\mu_{0}\right)$, so that the final result is given by,

$$
\begin{equation*}
C_{V}=\frac{1}{3} \pi^{2} k^{2} D\left(\mu_{0}\right) T \tag{0.23}
\end{equation*}
$$

$E+M \# 2$ (answer)


$$
{\underset{\hat{x}}{\hat{y}}}_{\hat{x}} \rightarrow \hat{z}
$$

a) intiny complex votativiu $1=\lambda R e^{-i \omega t} \delta(z-d) \hat{y}$

For z>d: $\underline{E}=\operatorname{Re}\left[\tilde{E}_{+} e^{i[b(z-d)-\omega t]}\right](-\hat{y})$

$$
\underline{B}=R_{e}\left[\tilde{E}_{+} e^{i[k(z-d)-\omega t]]}\right](+\hat{X})
$$

For $0<z<d$ :

$$
\begin{aligned}
E= & R\left[\hat{E}_{1} e^{-i[k(z-d)+\omega t]}\right][1-\hat{y} \mid \\
& +\operatorname{Re}\left[\sqrt{R} \tilde{E}_{-} e^{i k d} e^{i[k z-\omega t]}\right](+\hat{y}) \\
B= & \operatorname{Re}\left[\hat{E}_{-} e^{-i[k(z-d)+\omega t]}\right](-\hat{x}) \\
& +\operatorname{Re}\left[\sqrt{R} \tilde{E}_{-} e^{i k d} e^{i[k z-\omega t]}\right](-\hat{x})
\end{aligned}
$$

$R$-pepporer $\Rightarrow \sqrt{R}$ in cumplitude reflection corfficient
b) Power permint valumisis $\frac{d P}{d V}=E \cdot \underline{f}$
putegrate ouer shut and tini average

$$
\left\langle\frac{d}{d A} P_{\rangle}=\left\langle\int_{-\infty}^{\infty} d z E \cdot \frac{\lambda}{2}\left(e^{-i \omega t}+e^{-i \omega t}\right) \delta(z-d) \hat{y}\right\rangle\right.
$$

$\left\{\left\langle\frac{d}{d A}\right\rangle=-\frac{1}{2} \lambda R_{e}\left(\tilde{E}_{+}\right)\right\} \Rightarrow$ nud to determine $\tilde{E}_{+}$from Buundary conditiona

Frelhethie fuld : $\Delta E_{y}=0$ for all $t$ at $z=d$

$$
\begin{aligned}
\Rightarrow & -\tilde{E}_{+}=-\tilde{E}_{-}+\sqrt{R} E_{-} e^{i 2 k d} \\
& \Rightarrow \tilde{E}_{+}=\left(1-\sqrt{R} e^{i z d}\right) \tilde{E}
\end{aligned}
$$

For wapitic field: $\quad \Delta B_{x}=\frac{4 \pi}{c} \lambda$ for all $t$ at $z=d$

$$
\begin{aligned}
& \tilde{E}_{+}-\left[-\tilde{E}_{-}-\sqrt{R} e^{i 2 k d} \tilde{E}_{-}\right]=\frac{4 \pi}{c} \lambda \\
& \tilde{E}_{f}+\left[1+\sqrt{R} e^{i 2 h d}\right] \tilde{E}=\frac{4 \pi}{c} \lambda
\end{aligned}
$$

but

$$
\begin{gathered}
{\left[1-\sqrt{R} e^{i 2 k d}\right] \tilde{E}_{-}+\left[1+\sqrt{R} e^{i 2 k d}\right] \tilde{E}_{-}=\frac{4 \pi}{c} \lambda} \\
\Rightarrow \tilde{E}_{-}=\left(\frac{1}{2}\right)\left(\frac{4 \pi}{c} \lambda\right) \\
\Rightarrow E_{+}=\frac{2 \pi}{c} \lambda\left(1-\sqrt{R} e^{i 2 k d}\right)
\end{gathered}
$$

and Ne power meded to maintain unrent 10:

$$
\left\langle\frac{d P}{d R}\right)=-\frac{\pi \lambda^{2}}{C}(1-\sqrt{R} \cos 2 k d)
$$

4) Electron spin in mogn. field:

$$
\begin{aligned}
& \hat{H}=-\vec{\mu} \cdot \vec{B}=2 \mu_{B} B \hat{S}_{z}=\mu_{\beta} B \hat{\sigma}_{z} \\
& \Rightarrow \hat{H}=\mu_{B} B\left(\begin{array}{cc}
1 & 0 \\
0 & -1
\end{array}\right) \\
& t=0 \quad|\psi(0)\rangle=\frac{1}{\sqrt{2}}\binom{1}{1}=\frac{1}{\sqrt{2}}\binom{1}{0}+\frac{1}{\sqrt{2}}\binom{0}{1}
\end{aligned}
$$

written in terms of eigenvectors of $\hat{H}$ (eizenwolenes

$$
\Rightarrow|\psi(t)\rangle=\frac{1}{\sqrt{2}} e^{i \mu_{B} B t / \hbar}\binom{1}{0}+\frac{1}{\sqrt{2}} e^{-i \mu_{B} B t / \mu_{B}}\binom{0}{1}
$$

To colurlate $\left\langle\delta_{x}\right\rangle$, note that $\hat{\sigma}_{x}\binom{1}{0}=\binom{0}{1}$,

$$
\begin{aligned}
& \hat{\sigma}_{x}\binom{0}{1}=\binom{1}{0} \Rightarrow\left\langle\hat{\sigma}_{x}\right\rangle=\langle\psi(t)| \hat{\sigma}_{x}|\psi(t)\rangle \\
& =\frac{1}{2} e^{2 i \mu_{\beta} B t / t}+\frac{1}{2} e^{-2 i \rho_{\beta} \beta t / t}
\end{aligned}
$$

$$
\begin{aligned}
& \Rightarrow\left\langle\hat{\sigma}_{x}\right\rangle=\cos \left(\frac{2 \mu_{B} B}{t_{1}} t\right)=\cos (\omega t) \\
& \omega=\frac{2 \mu_{B} B}{t_{h}}
\end{aligned}
$$

Similarly, $\quad \hat{\sigma}_{y}\binom{1}{0}=\binom{0}{i}, \quad \hat{\sigma}_{y}\binom{0}{1}=\binom{-i}{0}$

$$
\text { and }\left\langle\hat{\sigma}_{y}\right\rangle=\frac{1}{2} i e^{2 i \beta_{\beta} B t / \hbar}-\frac{1}{2} i e^{-2 i \cdots}
$$

$$
=-\sin (\omega t)
$$

$\rightarrow$ the spin precesses with ong-prep. $\omega$ around the direction of $\vec{B}$ :
c)
 (clockwise looking down on $\vec{B}$, which is right since $q_{e}<0$ ).
perturbation

$$
E^{(1)}=\operatorname{re}_{e} B_{1}(1,0)\left(\begin{array}{ll}
0 & 1 \\
1 & 0
\end{array}\right)\binom{1}{0}=0 \text { for the }
$$

eigoner. $E^{(0)}=\mu_{\beta} B$ and some for the eigenr.

$$
E^{(0)}=-\mu_{B} B
$$

ie. first order corrections are zero.

$$
\begin{aligned}
& E^{(2)}= \pm \frac{1}{2 \mu_{\beta} B}\left|\hat{V}_{12}\right|^{2} \quad \text { ore the second ordo } \\
& \text { corrections fr r } \\
& E^{(0)}= \pm \mu_{B} B \\
& V_{12}=\mu_{B} B_{1}(1,0)\left(\begin{array}{ll}
0 & 1 \\
1 & 0
\end{array}\right)\binom{0}{1}=\mu_{B} B_{1} \\
& \Rightarrow E^{(2)}= \pm \frac{\left(\mu_{B} B_{1}\right)^{2}}{2 \mu_{B} B}= \pm \frac{1}{2} \mu_{\beta} B\left(\frac{B_{1}}{B}\right)^{2}
\end{aligned}
$$

The exact result is of cousin $E= \pm \mu_{B} B_{\text {tot }}$ where $B_{\text {tot }}=\sqrt{B^{2}+B_{1}^{2}}$
ont since $B_{\text {tot }}=B^{2} \sqrt{1+\left(B_{1} / B\right)^{2}} \simeq B^{2}\left[1+\frac{1}{2}\left(\frac{B_{1}}{B}\right)^{2}\right]$
to lowest roles you got $E^{(1)}= \pm \rho_{\beta} B \frac{1}{2}\left(\frac{B_{i}}{B}\right)^{2}$.

## 2012 Comprehensive Exam

0.4 QM1

Consider a particle subject to a one-dimensional simple harmonic oscillator potential, whose Hamiltonian is

$$
H=\frac{p^{2}}{2 m}+\frac{k x^{2}}{2}
$$

where $p=-i \hbar \partial_{x}$ is the particle's momentum operator, $m$ mass, and $k$ spring constant. Suppose that at $t=0$ the state vector is given by

$$
e^{-i p a / \hbar}|0\rangle,
$$

where $|0\rangle$ is the ground state and $a$ is some number with dimension of length. Evaluate the expectation value of position $\langle x\rangle$ for $t \geq 0$.

Solution: At $t=0$,

$$
\langle p\rangle_{0}=\langle 0| e^{i p a / \hbar} p e^{-i p a / \hbar}|0\rangle=\langle 0| p|0\rangle=0
$$

and

$$
\langle x\rangle_{0}=\langle 0| e^{i p a / \hbar} x e^{-i p a / \hbar}|0\rangle=\langle 0| e^{i p a / \hbar}\left[x, e^{-i p a / \hbar}\right]|0\rangle=i \hbar\langle 0| e^{i p a / \hbar} \partial_{p} e^{-i p a / \hbar}|0\rangle=a
$$

Using harmonic oscillator equations of motion,

$$
\dot{x}=\frac{i}{\hbar}[H, x]=\partial_{p} H=\frac{p}{m} \quad \text { and } \quad \dot{p}=\frac{i}{\hbar}[H, p]=-\partial_{x} H=-k x
$$

we immediately find the oscillatory solution in the form

$$
\langle x\rangle_{t}=a \cos (\omega t)
$$

were $\omega=\sqrt{k / m}$ is the oscillator natural frequency.

Quantum mechanics

1) A free, spinless non-relativistic particle with mass $m$ and charge $q$ is moving in a uniform magnetic field $\mathrm{B}=\mathrm{Bz}$. Find the spectrum of energy eigenvalues.
a) Write down the Hamiltonian H in terms of the conjugate momentum $\pi$. including the magnetic field via the minimal substitution $p \rightarrow \pi-p-q A$.
b) Choose a convenient gauge and write the vector potential A in terms of B .
c) Calculate $[\pi, \pi, \pi$.
d) Write $\pi_{5}=i c\left(a-a^{\dagger}\right)$ and $\pi_{y}=c\left(a+a^{\dagger}\right)$. where $c$ is a constant. $a$ is an annihilation operator, and $\mathbf{a}^{\dagger}$ is a creation operator. Find $c$.
c) Find the eigenvalues of H .
a) $H=\frac{\pi^{2}}{2 m}=\frac{(\vec{p}-9 \vec{b})^{2}}{2 m}$


$$
A=(0, \times B, 0)
$$

c)

$$
\begin{aligned}
{\left[\pi_{x}, \pi_{y}\right] } & =\left[p_{x}-q A_{x}, p_{y}-q A_{y}\right]=\left[p_{x}, p_{j}-q \times B\right] \\
& =-q B\left[p_{x}, x\right]=i \hbar q B
\end{aligned}
$$

d)

$$
\begin{aligned}
{\left[a, a^{+}\right]=1 \quad\left[\pi_{x}, \pi_{y}\right] } & =i c^{2}\left[a-a^{+}, a+a^{+}\right] \\
c= \pm \sqrt{\mathrm{hq} B} & =i c^{2}\left(\left[a, a^{+}\right]-\left[a^{+}, a\right]\right)=2 i c^{2}=i \hbar q B
\end{aligned}
$$

e)

$$
\begin{aligned}
& \left.A=\frac{\pi_{x}^{2}+\pi y^{2}+\pi z_{z}^{2}}{2 m}=\frac{1}{2 m}\left(-\frac{\hbar q B}{2}\left(a-a^{+}\right)\left(a-a^{+}\right)+\frac{\hbar q B}{2}(a+a)\right)\left(a+a^{+}\right)+P_{z}^{2}\right) \\
& =\frac{1}{2 m}\left(\frac{\tan \beta}{2}\right)\left(-a a+a^{\prime} a^{+}+a a^{+}+a^{1} a++a^{2}+d+a^{+}+a^{+} a+a a^{2}\right)+\frac{p_{2}^{2}}{2 m}
\end{aligned}
$$

## $Q .6$

A particle in a spherically symmetrical potential is known to be in an eigenstate of $\mathbf{L}^{2}$ and $L_{z}$ with eigenvalues $\hbar^{2} l(l+1)$ and $m \hbar$, respectively, denoted by $|l m\rangle . \mathbf{L}$ is the angular momentum operator, whose components obey the usual commutation algebra. Prove that the expectation values involving $L_{x}$ and $L_{y}$ obey

$$
\left\langle L_{x}\right\rangle=\left\langle L_{y}\right\rangle=0,\left\langle L_{x}^{2}\right\rangle=\left\langle L_{y}^{2}\right\rangle=\frac{l(l+1)-m^{2}}{2} \hbar^{2}
$$

in the eigenstate $|l m\rangle$.
Solution: Using $\left[L_{i}, L_{j}\right]=i \hbar \epsilon_{i j k} L_{k}$, we evaluate (henceforth setting $\hbar=1$ )

$$
\begin{aligned}
\left\langle L_{x}^{2}-L_{y}^{2}\right\rangle= & \left\langle\left[L_{x}, L_{z}\right]^{2}-\left[L_{y}, L_{z}\right]^{2}\right\rangle=2 m\left\langle L_{x} L_{z} L_{x}-L_{y} L_{z} L_{y}\right\rangle-m^{2}\left\langle L_{x}^{2}-L_{y}^{2}\right\rangle-\left\langle L_{x} L_{z}^{2} L_{x}-L_{y} L_{z}^{2} L_{y}\right\rangle \\
= & 2 m\left\langle L_{x} L_{z} L_{x}-L_{y} L_{z} L_{y}\right\rangle-m^{2}\left\langle L_{x}^{2}-L_{y}^{2}\right\rangle \\
& -\left\langle L_{x} L_{z}\left[L_{z}, L_{x}\right]+m L_{x} L_{z} L_{x}-\left[L_{y}, L_{z}\right] L_{z} L_{y}-m L_{y} L_{z} L_{y}\right\rangle \\
= & m\left\langle L_{x} L_{z} L_{x}-L_{y} L_{z} L_{y}\right\rangle-m^{2}\left\langle L_{x}^{2}-L_{y}^{2}\right\rangle \\
= & m\left\langle L_{x}\left[L_{z}, L_{x}\right]+m L_{x}^{2}-\left[L_{y}, L_{z}\right] L_{y}-m L_{y}^{2}\right\rangle-m^{2}\left\langle L_{x}^{2}-L_{y}^{2}\right\rangle=0 .
\end{aligned}
$$

Thus (restoring $\hbar$ )

$$
\left\langle L_{x}^{2}\right\rangle=\left\langle L_{y}^{2}\right\rangle=\frac{\left\langle\mathbf{L}^{2}-L_{z}^{2}\right\rangle}{2}=\frac{l(l+1)-m^{2}}{2} \hbar^{2} .
$$

The other identities are also easily obtained:

$$
i \hbar\left\langle L_{x}\right\rangle=\left\langle\left[L_{y}, L_{z}\right]\right\rangle=m\left\langle L_{y}-L_{y}\right\rangle=0
$$

and similarly for $\left\langle L_{y}\right\rangle$.

$$
H=\left(\begin{array}{ccc}
E_{1} & 0 & a \\
0 & E_{1} & b \\
a^{*} & b^{*} & E_{2}
\end{array}\right) .
$$

Let $H=H_{0}+V$ where

$$
V=\left(\begin{array}{ccc}
0 & 0 & a \\
0 & 0 & b \\
a^{*} & b^{*} & 0
\end{array}\right)
$$

(a.). Using non-degenerate penturbation Heory, if $E_{m}=E_{m}^{0}+\Delta_{m}^{\prime}+\Delta_{m}^{2}+\cdots$

$$
\Delta_{m}^{\prime}=\left\langle m_{0}\right| V\left|m_{0}\right\rangle
$$

For $\quad|1\rangle=,\left(\begin{array}{l}1 \\ 0 \\ 0\end{array}\right), \quad \Delta_{m}^{\prime}=V_{11}=0$.

$$
\begin{aligned}
& |2\rangle=\left(\begin{array}{l}
0 \\
1 \\
0
\end{array}\right), \quad \Delta_{m}^{2}=V_{22}=0 . \\
& |3\rangle=\left(\begin{array}{l}
0 \\
0 \\
1
\end{array}\right), \quad \Delta_{m}^{3}=V_{33}=0 .
\end{aligned}
$$

At second -ordes
(b).

$$
\begin{aligned}
& \operatorname{det}(H-\lambda I) . \\
& =\left(\begin{array}{ccc}
E_{1}-\lambda & 0 & a \\
0 & E_{1}-\lambda & b \\
a^{*} & b^{*} & E_{2}-\lambda
\end{array}\right) \\
& =\left(E_{1}-\lambda\right) \cdot\left\{\left(E_{1}-\lambda\right)\left(E_{2}-\lambda\right)-|b|^{2}\right\} \\
& -|a|^{2} \cdot\left(E_{1}-\lambda\right) \\
& =\left(E_{1}-\lambda\right)\left\{\lambda^{2}-\left(E_{1}+E_{2}\right) \lambda+E_{1} E_{2}\right. \\
& \left.-|b|^{2}-|a|^{2}\right\} .
\end{aligned}
$$

$$
\begin{aligned}
& \operatorname{det}(H-\lambda I)=0 \\
& \Rightarrow \lambda=E_{1} \quad \sigma \\
& \lambda=\frac{E_{1}+E_{2} \pm \sqrt{\left(E_{1}+E_{2}\right)^{2}-4 \cdot\left(E_{1} E_{2}-|b|^{2}-|a|^{2}\right)}}{2} \\
&=\frac{E_{1}+E_{2}}{2} \pm \sqrt{\left(\frac{E_{1}+E_{2}}{2}\right)^{2}-E_{1} E_{2}+|a|^{2}+|b|^{2}} \\
&=\frac{E_{1}+E_{2}}{2} \pm \sqrt{\left(\frac{E_{1}-E_{2}}{2}\right)^{2}+|a|^{2}+|b|^{2}}
\end{aligned}
$$

(c) Let $11^{\circ}>$, $12^{\circ}>$ be the appropriate
vases for degenerate perturbation theory.
for the degenerate sulespace of. $H_{0}$.

Then, to $i=1,2$

$$
\begin{aligned}
&\left(\mu_{0}-E_{1}\right) \cdot\left\{\left|i^{0}\right\rangle+\left|i^{\prime}\right\rangle+\cdots\right. \\
&=\left(\Delta_{i}^{\prime}+\Delta_{i}^{2} \cdots-V\right) \cdot\left\{\left|i^{\prime 0}\right\rangle+\left|i^{\prime}\right\rangle+\cdots\right. \\
&-A
\end{aligned} .
$$

Operate on L.H.S 2 R.H.S. with $\langle j 01$ where $j \in\{1,2\}$.
we find.

$$
\begin{align*}
& \text { find } \\
& \text { L.H.S }=0=\left(\Delta_{i}^{\prime}\left\langle j 0 \mid i^{0}\right\rangle-\langle j| V\left|i^{0}\right\rangle .\right) \\
&+\left(\Delta_{i}{ }^{2}\left\langle j \mid i^{0}\right\rangle-\left\langle j^{0}\right| V\left|i^{\prime}\right\rangle\right)  \tag{B}\\
&+\cdots
\end{align*}
$$

where we have grouped terns of the
same order together.

Operate on (A) with $\left\langle 3^{\circ}\right\rangle=\langle 31$, we
find.

$$
\begin{aligned}
& \left(E_{2}-E_{1}\right) \cdot\left\{\left\langle 3^{0} \mid i^{0}\right\rangle+\left\langle 3^{0} \mid i^{\prime}\right\rangle+\cdots\right\} \\
= & \Lambda_{i}^{1}\left\langle 3^{0} \mid i^{0}\right\rangle-\left\langle 3^{0}\right| . V\left|i^{0}\right\rangle
\end{aligned}
$$

+ higher. order terms.

From (C), we see comparing $1^{\text {st }}$ order terms,

$$
\left\langle 3^{0} \mid i^{\prime}\right\rangle=\frac{\left\langle 3^{0}\right| V\left|i^{0}\right\rangle}{E_{1}-E_{2}}
$$

From (B), $1^{\text {st }}$ order terms give us:
$\langle j| V\left|i^{0}\right\rangle=0$. wot this is true for any states in the degenerate space spanned by 11$\rangle \quad 212\rangle$.

At $2^{\text {rd }}$ order, we get

$$
\Delta^{2} \delta_{j_{0}, i_{0}}=\left\langle j^{0}\right| \vee\left|i^{i}\right\rangle
$$

In other words. for the right vases

$$
\left\langle j^{0}\right| \vee\left|.3^{0}\right\rangle\left\langle 3^{0}\right| \vee\left|i^{0}\right\rangle=\Delta^{2} \delta_{i_{0}, j 0}
$$

Thus we have to diagonalize the matrix, $M=\frac{P_{D} \vee\left|3^{0}\right\rangle\left\langle 3^{0}\right| V P_{D}}{E_{1}-E_{2}}$. where

$$
P_{D}=|1\rangle\langle 1|+|2\rangle\langle 2| .
$$

In the basis $,|1\rangle=\left(\begin{array}{l}1 \\ 0 \\ 0\end{array}\right), \quad|2\rangle=\left(\begin{array}{l}0 \\ 1 \\ 0\end{array}\right)$,

$$
\begin{aligned}
& M_{11}=\frac{V_{13} V_{31}}{E_{1}-E_{2}}=\frac{|a|^{2}}{E_{1}-E_{2} .} \\
& M_{12}=\frac{V_{13} V_{32}}{E_{1}-E_{2}}=\frac{a b^{*}}{E_{1}-E_{2}} \\
& M_{21}=\frac{a^{*} b}{E_{1}-E_{2}} . \quad M_{22}=\frac{|b|^{2}}{E_{1}-E_{2}} .
\end{aligned}
$$

To diagonalize $M$,

$$
\begin{aligned}
& \operatorname{det}(M-\lambda I)=0 \cdot \\
& \Leftrightarrow \quad \operatorname{det}\left(\begin{array}{ccc}
|a|^{2}-\lambda & a b^{*} \\
a^{*} & b & |b|^{2}-\lambda
\end{array}\right)=0 . \\
& \Leftrightarrow \quad\left(|a|^{2}-\lambda\right)\left(|b|^{2}-\lambda\right)-|a|^{2}|b|^{2}=0 \\
& \Leftrightarrow \quad \lambda^{2}-\lambda\left(|a|^{2}+|b|^{2}\right)=0 \\
& \Leftrightarrow \quad \lambda=0 \quad a \quad \lambda=|a|^{2}+|b|^{2} .
\end{aligned}
$$

thus the eigenvalues of $M$ are

$$
0, \quad \frac{|a|^{2}+|b|^{2}}{E_{1}-E_{2}}
$$

Fo $\quad|3\rangle$,

$$
\Delta_{3}^{(2)}=\frac{\left.\left|\left\langle 1^{\circ}\right| V\right| 3^{\circ}\right\rangle\left.\right|^{2}+\left|\left\langle 2^{0} \mid V / 3^{\circ}\right\rangle\right|^{2}}{E_{2}-E_{1} .} .
$$

which gives the same result as non-degenenats perturbation theory,

$$
\Delta_{3}^{(2)}=-\frac{\left(|a|^{2}+|b|^{2}\right)}{E_{1}-E_{2}}
$$

from the exact results,

$$
\begin{aligned}
E & =E_{1} \\
& =\frac{E_{1}+E_{2}}{2} \pm \sqrt{\frac{\left(E_{1}-E_{2}\right)^{2}}{2}+|a|^{2}+|b|^{2}} \\
& =\frac{E_{1}+E_{2}}{2} \pm \frac{E_{1}-E_{2}\left(1+\frac{1}{2}\left(|a|^{2}+|b|^{2}\right) \frac{12}{\left(E_{1}-E_{2}\right)^{2}}+\cdots\right.}{}= \\
& E_{1}+\frac{\left(|a|^{2}+|b|^{2}\right)}{E_{1}-E_{2}}+\cdots \\
& E_{2}-\frac{\left(|a|^{2}+|b|^{2}\right)}{E_{1}-E_{2}}
\end{aligned}
$$

which agrees with the results of
degemente perturbation the org, bout
not with those of non-degenerate serturtation theory, as expected.

Q8

Statistical mechanics
2) This problem concerns the fundamental definitions of thermodynamics. The numbers hate been kept small to minimize the mathematics - use discrete differences in place of derivatives where appropriate.

A system has $\mathrm{N}=2$ distinguishable particles, each of which can have energy $0, \varepsilon, 2 \varepsilon, 3 \varepsilon$. Doc. Say the system has total energy $E=5 \varepsilon$.
a) What is the entropy $S$ ?
b) What is the temperature $T$ ?
c) What is the chemical potential $\mu$ ? (Hint: add a $3^{\text {rd }}$ particle.)

Imagine a different system with $N=6$ distinguishable particles. each of which can have energy 0 or $\varepsilon$.
d) Plot the entropy as a function of the different discrete values of the total $E$.
e) For what value of $E$ is the temperature maximized?
f) Plot $T$ as a function of $E$ for the case where $N$ is large (but fixed), and comment on the various limiting cases.

$$
\begin{aligned}
& \frac{\varepsilon_{1} \varepsilon_{2}}{0 \quad 5 \varepsilon_{1}} \\
& \text { There are } 6 \text { accessible states } \Omega \\
& \text { a) } S=k_{B} \ln 6 \\
& \begin{array}{ll}
1 & 4 \varepsilon \\
2 & 3 \varepsilon \\
3 & 2 \varepsilon \\
4 & 1 \varepsilon
\end{array} \\
& \frac{1}{T}=\frac{\partial S}{\partial E}=\frac{\Delta S}{\Delta E}=\frac{S_{6 \varepsilon}-S_{S \varepsilon}}{6 \varepsilon-S \varepsilon}=\frac{k_{B} \ln \left(\frac{6}{5}\right)}{\varepsilon} \\
& 50 \varepsilon \\
& \text { b) } \left.T=\frac{q}{k_{3} \ln \left(\frac{6}{5}\right)}\right]
\end{aligned}
$$

We require the entropy to remain constant as we add the 3nd particle.
 1 total energy of $2 \varepsilon$ gives 6
accessible states accessible states, as before.

| 0 | 2 | 0 |
| :--- | :--- | :--- |
| 0 | 0 | 2 |
| 1 | 1 | 0 |
| 1 | 0 | 1 |
| 0 | 1 | 1 |

$$
\mu=\frac{2 \varepsilon-5 \varepsilon}{3 \text { pinks }{ }^{2} \text { pmidicks }}=-3 \varepsilon=\mu
$$

d) $M=6$, each with $\varepsilon=0$ on $\varepsilon$
e) The temperature is maximized for $E=2 \varepsilon$. ( $T=\frac{\partial S}{\partial E}$ is undefined at $E=3 \varepsilon$ )
f)


For $E \rightarrow 0$ or $N \varepsilon$
He temperature $\rightarrow 0$ from positive on negative values respectively.

For $E \rightarrow \frac{N}{2} \varepsilon, T \rightarrow \pm \infty$, depending on whether he approach is from above or below. For $E>\frac{N}{2} \varepsilon$ the temperatures are negative.

Capillary waves

$$
\omega=\alpha h^{3 / 2}, \alpha=\sqrt{\frac{\sigma}{\rho}} \text { disp.rel. }
$$

Bosons, $\mu=0$ (number of particles not weasand)
$\rightarrow$ occupation numans $\quad M_{k}=\frac{1}{e^{E_{k} / T}-1}$

$$
\begin{aligned}
& E_{k}=\text { hm in in: } \\
& E=\int \frac{d^{2} p d^{2} x}{h^{2}} \frac{\pi \omega}{e^{\hbar \omega / T}-1} \quad \text { where } p=t_{k} k \\
& \Rightarrow E=\frac{A}{h^{2}} \int_{0}^{\infty} 2 \pi p d p \frac{\hbar \omega}{e^{\frac{\pi}{\hbar} \omega / T}-1} \quad A \text { is the area. }
\end{aligned}
$$

Change rariobl, to $\omega: \quad P=\pi k$,

$$
\begin{aligned}
& d \omega=\frac{3}{2} \alpha k^{1 / 2} d h \quad, k^{1 / 2}=\left(\frac{\omega}{\alpha}\right)^{1 / 3} \\
& \Rightarrow E=\frac{2 \pi A}{h^{2}} \hbar^{2} \int_{0}^{\infty} d \omega \frac{2}{3 \alpha}\left(\frac{\omega}{\alpha}\right)^{1 / 3} \frac{\hbar^{2} \omega}{e^{\pi \omega / T}-1} \\
& \frac{\pi \omega}{T}=x
\end{aligned}
$$

So $E \propto T^{7 / 3}$ and $C_{A}=\left.\frac{\partial E}{\partial T}\right|_{A} \propto T^{4 / 3}$

Q 10.
SM \#1 (answer)

where $B=\nabla \times A$
Hanittorian for maguatized particle:

$$
H=\frac{1}{2 m}\left(p-\frac{q}{c} A\right)^{2}=\frac{1}{2 m}\left[\left(P_{x}-\frac{-\mu s}{2}\right)^{2}+\left(P_{y}+\frac{q B_{0}}{2 c} x\right)^{2}+P_{z}^{2}\right]
$$

N-Classical, zon-interacting porticles howi a partifun furetive : $z_{N}=\frac{1}{N!}\left[\int \frac{d p_{x} d p_{y} d p_{z} d x d y d z e^{-\beta H}}{h^{3}}\right]^{N}$
 curider:

$$
\begin{aligned}
& \text { corridir: } \\
& {\left[\int_{-\infty}^{\infty} d p_{x} \int_{-\infty}^{\infty} d p_{y} \int_{-\infty}^{\infty} d p_{z} \int d x y d z e^{-\frac{\beta}{2 m}\left[\left(p_{x}-\frac{9 y \sigma_{0}}{2 c}\right)^{2}+\left(p_{y}+\frac{q x \beta_{0}}{2}\right)^{2}+p_{z}^{2}\right]}\right.}
\end{aligned}
$$

chang variables to: $u_{x} \equiv P_{x}-\frac{q y_{0}}{2 c} ; u_{y}=p_{y}+\frac{q x B_{0}}{2 c}$

$$
\begin{aligned}
& \rightarrow \int_{-\infty}^{\infty} d u_{x} \int_{-\infty}^{\infty} d u_{y} \int_{-\infty}^{\infty} d q_{z} \int d x d y d z e^{e^{-\frac{\beta}{2 m}\left[u_{x}^{2}+u_{y}^{2}+u_{z}^{2}\right]}} \\
& \qquad V_{[-\infty}\left[\int_{-\infty}^{-} d u_{x} e^{-\frac{p}{2 m} u_{x}^{2}}\right]^{3}=v(2 \pi m \cdot k T)^{3 / 2}
\end{aligned}
$$

$$
\Rightarrow z_{N}=\frac{1}{N!} \frac{(2 \pi m k T)^{3 N / 2} V^{N}}{h^{3 N}}
$$

is midependent of $B_{0} \Rightarrow$ diku a anual
umagnitioet chal ger
m. Q10, page2

SM \#1 (Guswer ...)
a) Eq.of state is: $\quad \bar{p}=\frac{1}{\beta} \frac{\partial}{\partial v} \ln z_{N}=\frac{1}{\beta} \frac{\partial}{\partial v}[N \ln v]$

$$
\Rightarrow \quad \bar{p}=\frac{N k T}{V} \text { idhalgas } E_{q} \cdot \text { of state }
$$

b) $\left\{\begin{array}{l}\left.\langle M\rangle=\frac{1}{\beta} \frac{\partial}{\partial B_{0}} \ln z_{N}=-0\right\} \text { which 's } \\ \text { Bohr's Thesis result. }\end{array}\right.$
c)

$$
\begin{aligned}
& \left\langle E^{2}\right\rangle-\bar{E}^{2}=\frac{\partial^{2}}{\partial \beta^{2}} \ln Z_{N}=-\frac{\partial}{\partial \beta} \bar{E} \\
& \text { with } \bar{E} \equiv\langle E\rangle=-\frac{\partial}{\partial \beta} \ln Z_{N}=-\frac{\partial}{\partial \beta}\left[1-\frac{3 N}{2} \ln \beta\right] \\
& \overline{\bar{E}}=\frac{3 N}{2} \frac{1}{\beta} \\
& \Rightarrow \quad \frac{\partial}{\partial \beta} \bar{E}=-\frac{3 N}{2 \beta^{2}}=-\frac{1}{\beta} \bar{E}=-\frac{2}{3 N} \bar{E}^{2} \\
& \Rightarrow \quad\left\langle E^{2}\right\rangle-\bar{E}^{2}=\frac{2}{3 N} E^{2} \\
& \Rightarrow \quad \sqrt{\frac{\left\langle E^{2}\right\rangle-E^{2}}{E^{2}}}=\sqrt{\frac{2}{3 N}}
\end{aligned}
$$

Probll.
principle of supecposition $\quad \underset{\sim}{H}={\underset{\sim}{2}}_{2}-H_{1}$
$\underline{H}_{2}=$ foid of solod, $\underline{H}_{1}=$ hild of hole
$I_{2}, I_{1}, \underline{i}$
$a:=\frac{D}{2}$ radius of small hale

$$
\begin{gathered}
I=I_{2}-I_{1}=9 \pi a^{2} j-\pi a^{2} j=8 \pi a^{2} j \\
\Rightarrow j=\frac{I}{8 \pi a^{2}} \\
I_{1}=\pi a^{2} j=\frac{I}{8} \\
I_{2}=9 \pi a^{2} j=\frac{9}{8} I
\end{gathered}
$$

cunced flows aul of pape plaine
Ampere's law

$$
\begin{aligned}
& H_{2 x}=-\frac{I y}{16 \pi a^{2}}, H_{2 y}=\frac{1 x}{16 \pi a^{2}} \quad \begin{array}{l}
\text { inside } \\
(r \leq 3 a
\end{array} \\
& \text { ( } \left.r=\sqrt{x^{2}+y^{2}}\right) \\
& H_{2} x=-\frac{9 I_{y}}{16 \pi\left(x^{2}+y^{2}\right)}, H_{2 y}=\frac{9 I_{x}}{16 \pi\left(x^{2}+y^{2}\right)} . \\
& \text { ontside } \\
& (r>3 a)
\end{aligned}
$$

$$
\begin{equation*}
H_{1 x}=-\frac{I y}{16 \pi a^{2}}, H_{y y}=\frac{I(x-a)}{16 \pi a^{2}} \quad\left(r_{1} \leq a\right) \tag{2}
\end{equation*}
$$

wee $\quad r_{1}=\sqrt{(x-a)^{2}+y^{2}}$ casions

$$
H_{1 x}=-\frac{I y}{16 \pi\left[(x-a)^{2}+y^{2}\right]}, H_{1 y}=\frac{I(x-a)}{16 \pi\left[(x-a)^{2}+y^{2}\right]} \quad(r,>a)
$$

On the plane $P, H_{2 x}=H_{1 x}=O$

$$
\begin{aligned}
\Rightarrow H_{x} & =0 \\
H_{y} & =H_{2 y}-H_{1 y}
\end{aligned}
$$

$\Rightarrow$ inside hole $(0<x<2 a)$

$$
H_{y}=\frac{I x}{16 \pi a^{2}}-\frac{I(x-a)}{16 \pi a^{2}}=\frac{I a}{16 \pi a^{2}}=\frac{I}{16 \pi a}
$$

inside solid $(2 a \leq x \leq 3 a$ or $-3 a \leq x \leq 0)$

$$
H_{y}=\frac{1 x}{16 \pi a^{2}}-\frac{I(x-a)}{16 \pi\left[(x-a)^{2}+y^{2}\right]}=\frac{I\left(x^{2}-a x-a^{2}\right)}{16 \pi a^{2}(x-a)}
$$

outside cylinder

$$
H_{y}=\frac{9 I x}{16 \pi\left(x^{2}+y^{2}\right)}-\frac{I(x-a)}{16 \pi\left[(x-a)^{2}+y^{2}\right]}=\frac{(8 x-9 a) I}{16 \pi x(x-a)}
$$

The magrelic held at ald points inside the hole

$$
\begin{aligned}
& \left(r_{1} \leq a\right) \\
& H_{x}=-\frac{I y}{16 \pi a^{2}}+\frac{I y}{16 \pi a^{2}}=0 \\
& H_{y}=\frac{1 x}{16 \pi a^{2}}-\frac{I(x-a)}{16 \pi a^{2}}=\frac{I}{16 \pi a}
\end{aligned}
$$

$\Rightarrow$ uniform and along y-axis

## Q. 12

A pulsar emits bursts of radio waves which are observed from the Earth at two different frequencies, $\omega_{1}$ and $\omega_{2}$. An astronomer notes that the arrival time of the bursts is delayed at the lower frequency, that is, the pulse at $\omega_{1}$ arrives after the pulse at $\omega_{2}$. The delay, $\tau$ is due to dispersion in the interstellar medium. Assume this medium consists of ionized hydrogen (called the dilute plasma), $m_{e}$ is the mass of the electron and $N$ the number of electrons per unit volume.
(a). Find the index of refraction of the dilute plasma.
(b). Find the distance from the pulsar to the Earth.

## Solution

(a) For an electron in the dilute plasma, we have $m_{e} \ddot{\vec{x}}=-e \vec{E}$ and $\vec{E}=\vec{E}_{0} e^{-i o t}$

Solving the above equaiton, we get the dipole moment of the electron,
$\vec{p}=-e \vec{x}=-\frac{\left(e^{2} / m\right)}{\omega^{2}} \vec{E}$
The polarization is given by
$\vec{P}=-\frac{\left(N e^{2} / m\right)}{\omega^{2}} \vec{E}$
As $\vec{P}=\varepsilon_{0} \chi_{e} \vec{E}$, we have $\varepsilon_{r}=1+\chi_{e}=1-\frac{\omega_{p}^{2}}{\omega^{2}}$ where $\omega_{p}=e \sqrt{\frac{N}{m \varepsilon_{0}}}$.
The index of refraction of the dilute plasma is $n=\sqrt{\varepsilon_{r}}=\sqrt{1-\frac{\omega_{p}^{2}}{\omega^{2}}}$
(b) The pulse the pulsar emits is a wave packet that travels to the Earth. As such, it travels at the group velocity, not the phase velocity.
$k=\frac{\omega}{c} n=\frac{\omega}{c} \sqrt{1-\frac{\omega_{p}^{2}}{\omega^{2}}}$
$\frac{1}{v_{g}}=\frac{\partial k}{\partial \omega}=\frac{1}{c}\left[\frac{\omega}{2}\left(1-\frac{\omega_{p}^{2}}{\omega^{2}}\right)^{-1 / 2}\left(\frac{2 \omega_{p}^{2}}{\omega^{3}}\right)+\left(1-\frac{\omega_{p}^{2}}{\omega^{2}}\right)^{1 / 2}\right]$
$=\cdots=\frac{1}{c}\left(1-\frac{\omega_{p}^{2}}{\omega^{2}}\right)^{-1 / 2}$
$\tau=t_{1}-t_{2}=l\left(\frac{1}{v_{g 1}}-\frac{1}{v_{g 2}}\right)=\frac{l \omega_{p}^{2}}{2 c}\left(\frac{1}{\omega_{1}^{2}}-\frac{1}{\omega_{2}^{2}}\right)$ where $l$ is the distance from the pulsar to the
Earth.

### 2.13

(a) Consider two positrons in a beam at SLAC. The beam has energy of about $50 \mathrm{GeV}(\gamma$ $\approx 10^{5}$ ). In the beam (rest) frame, they are separated by a distance $d$, and positron $e_{2}^{+}$is traveling directly ahead of $e_{1}^{+}$in the Z-axis, shown in the figure (left) below. Write down $\vec{E}, \vec{B}$, the Lorentz force $\vec{F}$ and the acceleration $\vec{a}$ on $e_{1}^{+}$exerted by $e_{2}^{+}$. Do this in both the rest and laboratory frames.
(b) The problem is the same as in part (a) except this time the two positrons are traveling side by side as shown in the figure (right) below.


Two positrons separated by a distance of $d$ travel with a velocity of $v$ in the $Z$ axis.

## Solution

(a) Let $\mathrm{K}^{\prime}$ and K the beam rest and laboratory frames, respectively. In frame K',
$\vec{E}^{\prime}=-\frac{1}{4 \pi \varepsilon_{0}} \frac{e}{d^{2}} \hat{z} \quad \vec{B}^{\prime}=0 \quad \vec{F}^{\prime}=-\frac{1}{4 \pi \varepsilon_{0}} \frac{e^{2}}{d^{2}} \hat{z} \quad \vec{a}^{\prime}=-\frac{1}{4 \pi m \varepsilon_{0}} \frac{e^{2}}{d^{2}} \hat{z}$
In frame $K$,
$\vec{E}=\vec{E}^{\prime}=-\frac{1}{4 \pi \varepsilon_{0}} \frac{e}{d^{2}} \hat{z} \quad \vec{B}=0 \quad \vec{F}=e \vec{E}=\vec{F}^{\prime}$
$\vec{a}=c \frac{d \beta}{d t}=\frac{1}{m \gamma}\left(F-F \beta^{2}\right)=\frac{F}{m \gamma^{3}}=\frac{\vec{a}^{\prime}}{\gamma^{3}}$
In this case, the EM filed and the Lorentz force are the same in K and $\mathrm{K}^{\prime}$. Due to the relativistic effect the acceleration of $e_{1}^{+}$in frame K is only $\frac{1}{\gamma^{3}}$ times that in the rest frame.
In other words, the force exerted by a neighboring collinear charge on a charge moving with high speed will be small.
(b) In frame $\mathrm{K}^{\prime}$,

$$
\vec{E}^{\prime}=-\frac{1}{4 \pi \varepsilon_{0}} \frac{e}{d^{2}} \hat{x} \quad \vec{B}^{\prime}=0 \quad \vec{F}^{\prime}=-\frac{1}{4 \pi \varepsilon_{0}} \frac{e^{2}}{d^{2}} \hat{x} \quad \vec{a}^{\prime}=-\frac{1}{4 \pi m \varepsilon_{0}} \frac{e^{2}}{d^{2}} \hat{x}
$$

In frame K ,
$\vec{E}=\gamma \vec{E}^{\prime} \quad \vec{B}=\gamma \frac{v}{c^{2}} E^{\prime} \hat{y} \quad \vec{F}=\frac{\vec{F}^{\prime}}{\gamma}$

$$
Q 13, p^{2}
$$

$$
\vec{a}=\frac{F}{m \gamma}=\frac{\vec{a}^{\prime}}{\gamma^{2}}
$$

Q 14.
$\frac{E+M \# 1(a n s w e r) .}{\hat{z}+}$
a)


$$
\nabla \cdot E=4 \pi\left(\rho_{\text {g人t }}+\rho_{p}\right) \mu \text { Pouwn's } \rho_{\rho}
$$

whare $\rho_{2 x}=q \delta(I) ; \rho_{p}$ patongene wind ing bsping displacument vestor $\bar{D}=\vec{G} \cdot E \quad T$ rampurms Pussuris Eq unto: $\nabla \cdot \underline{D}=4 \pi-q-\delta(\underline{r})$
sovalli by Gaure' law to: $D=\frac{9}{r^{2}} \hat{r}$
with $r=\left(x^{2}+y^{2}+z^{2}\right) ; \quad \hat{r}=\sin \theta \cos \phi \hat{x}+\sin \theta \sin \phi \xi^{1}+\operatorname{se\theta } \theta z$ an spharical coordemater

or $\hat{\gamma}=\frac{x}{r} \hat{x}+\frac{y}{T} \hat{y}+\frac{z}{r} \hat{z}$

$$
\begin{aligned}
& \Rightarrow \quad D=\frac{q}{r^{3}}[x \hat{x}+y \hat{y}+z \hat{z}] \\
& H \quad \stackrel{\epsilon}{\epsilon} \cdot \underline{\epsilon_{1}} E_{x} \hat{x}+\epsilon_{\perp} E_{y} \hat{y}+\epsilon_{11} E_{z} \hat{z} \\
& \Rightarrow E_{x}=\frac{q}{r^{3} \epsilon_{1}}, E_{y}=\frac{q}{r^{3} \epsilon_{1}} y, E_{z}=\frac{q}{r^{3} \epsilon_{11}}
\end{aligned}
$$

Tha magnitude is: $|E|=\sqrt{E_{x}^{2}+E_{y}^{2}+E_{z}^{2}}$

$$
\left\{|E|=\frac{9}{r^{3}}\left[\frac{x^{2}+y^{2}}{\epsilon_{+}^{2}}+\frac{z^{2}}{\epsilon_{11}^{2}}\right]^{1 / 2}\right\}
$$

EHM \#1 (answer/conl..)
b) For $r \neq 0$ all The charge deviate as due to polarization of The dilentic, $i \cdot E \quad \nabla \cdot E=4 \pi \rho_{p} \Rightarrow \rho_{p}=\frac{1}{4 \pi} \nabla \cdot E$ but $\nabla \cdot E=\partial_{x} E_{x}+\partial_{y} E_{y}+\partial_{z} E_{z}$
and $\partial_{x} E_{x}=\frac{q}{\epsilon_{I}} \frac{\partial}{\partial x}\left(\frac{x}{r^{3}}\right)=\frac{9}{\epsilon_{I}}\left[\frac{1}{r^{3}}-\frac{3 x^{2}}{r^{-54}}\right]$

$$
\begin{aligned}
& \partial_{y} E_{y}=\frac{9}{\epsilon_{1}} \frac{\partial}{\partial y}\left(\frac{y}{r^{2}}\right)=\frac{9}{\epsilon_{1}}\left[\frac{1}{r^{3}}-\frac{3 y^{2}}{r^{5}}\right] \\
& \Rightarrow \nabla \cdot E=\frac{9}{r^{5}}\left\{\frac{2\left(2 z^{2}-x^{2}-y^{2}\right)}{\epsilon_{1}}+\frac{x^{2}+y^{2}-2 z^{2}}{\epsilon_{11}}\right\} \\
& \left.\epsilon_{1}=\frac{\partial}{r^{3}}\right)=\frac{9}{\epsilon_{1}}\left[\frac{1}{r^{5}}\right] \\
& \nabla \cdot E=\frac{9}{r^{5}}\left[2 z^{2}-\left(x^{2}+y^{2}\right)\right]\left(\frac{1}{\epsilon_{1}}-\frac{1}{\epsilon_{11}}\right] \\
& \Rightarrow\left\{P_{p}(x, y, z)=\frac{9}{4 \pi r^{5}}\left[2 z^{2}-\left(x^{2}+y^{2}\right)\right]\left(\frac{1}{\epsilon_{1}}-\frac{1}{\epsilon_{11}}\right)\right.
\end{aligned}
$$

shows charge cones

c) Total elatrival euryy density $M u_{E}=\frac{1}{2} E \cdot D=\frac{1}{2} E \cdot E \cdot E$. but - $\stackrel{\rightharpoonup}{\epsilon} \cdot E=\epsilon_{I} E_{\bar{x}} \hat{x}+\epsilon_{1} E_{y} \hat{y}+\epsilon_{\| E} E_{z} \hat{z}$
$\downarrow E \cdot \vec{\epsilon}, \underline{E}=E_{1} E_{x}^{2}+e_{1} E_{y}^{2}+E_{11} E_{z}^{2}$

展
$u_{E}=\frac{1}{2}\left[\epsilon_{1} \frac{q^{2}}{r^{6} \epsilon_{I}^{2}} x^{2}+\frac{q^{2}}{r^{6} \epsilon_{1}^{2} y^{2}}+\frac{q^{2}}{r^{6} \epsilon_{11}^{2}} z^{z^{2}}\left[\frac{x^{2}+y^{2}}{\varepsilon_{1}}+\frac{z^{2}}{\epsilon_{11}}\right]\right]$

